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Instytut Matematyki, Uniwersytet Marii Curie-Skłodowskiej, Lublin Department of Mathematical Sciences, University of Delaware, Newark, Delaware, USA

Jan G. KRZYŻ, Richard J. LIBERA,⁴⁾ Eligiusz ZŁOTKIEWICZ

Coefficients of Inverses of Regular Starlike Functions

Współczynniki funkcji odwrotnych do funkcji regularnych gwiaździstych Коэффициенты функций обратных к регулярным звездным функциям

1. INTRODUCTION

As is usually the case we let \checkmark represent the class of functions of the form

(1.1)
$$f(z) = z + a_2 z^2 + a_3 z^3 + \cdots$$

regular and univalent in the open unit disk $\Delta = \{z \in C : |z| < 1\}$ Much of the interest in and many investigations of \mathscr{S} relate to establishing correct bounds on the coefficients a_k , $k = 2,3,\ldots$, and it has been shown, cf.e.g. [2], that $|a_n| \leq n$, for n = 2,3,4,5,6. Except for rotations the unique extremal for these bounds is the Koebe function

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(1.2)
$$k(z) = \frac{z}{(1-z)^2} = z + 2z^2 + 3z^3 + \cdots$$

In his seminal work relating to the conclusion that $|a_3| \leq 3$, Loewner [7] was able to give sharp bounds for the coefficients which appear in the Maclaurin series for the inverse of any function in S. Specifically, if the inverse of f(z) is

(1.3)
$$F(w) = w + \gamma_2 w^2 + \gamma_3 w^3 + \cdots$$

he showed that

(1.4)
$$|\gamma_n| \leq \frac{1}{n} \binom{2n}{n+1}$$

for $n \ge 2$ and that the sharp upper bound is achieved by the inverse of a rotation of k(z) defined by (1.2).

To summarize the situation briefly we can say that sharp bounds for $|\gamma_n|$ and each index n have been obtained in a surprisingly straightforward way, whereas proper bounds on $|a_n|$ have been obtained for only a few indices with great difficulty. The purpose of this note is to illustrate that the converse situation appears to hold for some well-known subclasses of S.

2. CONCLUSIONS

For $0 \leq \alpha \leq 1$ we let \mathcal{A}_{α}^{*} be the subclass of \mathcal{A} consisting of functions which are α -starlike, i.e., f(z) is as in (1.1) and $\operatorname{Re}\left\{zf'(z)/f(z)\right\} \geq \alpha$ for z in Δ . The functions f(z) in \mathcal{A} for which $f[\Delta]$, the image of Δ under f(z), is a convex domain is denoted by K; it is

well-known that K<S.

The family of all starlike functions is \mathcal{S}_{o}^{*} , written simply as \mathcal{S}^{*} .

Also, let P be the class of functions

(2.1)
$$P(z) = 1 + p_1 z + p_2 z^2 + \cdots$$

regular and satisfying the condition Re P(z) > 0 for z in \triangle . It follows that f(z) is in \triangle_{∞}^{*} if and only if there is a corresponding function P(z) in p for which

(2.2)
$$zf'(z) = f(z) (1 - \alpha)P(z) + \alpha$$

With representations (1.1) and (2.1) the last relation yields the relationships

(2.3)
$$(n-1)a_n = (1-\alpha)\sum_{j=1}^{n-1} p_j a_{n-j-1}, \quad n = 2,3,...$$

Now, if a function and its inverse are given by (1.2) and (1.3) a brief calculation shows that

(2.4)
$$\gamma_2 = -a_2$$
, $\gamma_3 = 2a_2^2 - a_3$ and $\gamma_4 = 5a_2[a_3 - a_2^2] - a_4$
and these along with (2.3) give $\gamma_2 = -(1 - \alpha)p_4$ and

(2.5)
$$\gamma_3 = -(\frac{1-\alpha}{2})[p_2 - 3(1-\alpha)p_1^2],$$

which give rise to the following result.

THEOREM 1. If f(z) is in S_{∞}^{*} and its inverse is given by (1.3), then $|\gamma_{2}| \leq 2(1 - \alpha)$ and

(2.6)
$$|\gamma_3| \in \begin{bmatrix} (1-\alpha)(5-6\alpha) & \underline{for} & 0 \le \alpha \le \frac{2}{3} \\ (1-\alpha) & \underline{for} & \frac{2}{3} \le \alpha < 1. \end{bmatrix}$$

106 J.G. Krzyż, R. J. Libera, E. J. Złotkiewicz These bounds are sharp.

The first bound follows from the relation $|p_k| \leq 2$ which is valid for all coefficients of (2.1) and the second is a consequence of the following lemma which is quoted in [6].

LEMMA. If P(z) in P is given by (2.1), then

(2.7)
$$|p_2 - \mu p_1^2| \le 2 \max\{1, |1 - 2\mu|\}$$

and the bound is rendered sharp by Q(z) = (1 + z)/(1 - z)then $|1 - 2\mu| \ge 1$ and by $T(z) = (1 + z^2)/(1 - z^2)$ otherwise.

Now, replacing P(z) in (2.2) by Q(z) and T(z) and solving for the corresponding f(z) gives functions in \mathscr{I}_{∞}^{*} , namely

(2.8)
$$k_{\alpha}(z) = \frac{z}{(1-z)^{2}(1-\alpha)} = z + 2(1-\alpha)z^{2} + (1-\alpha)(3-2\alpha)z^{3} + \dots$$

and

(2.9)
$$h_{\alpha}(z) = \frac{z}{(1-z^2)^{1-\alpha}} = z + (1-\alpha)z^3 - \frac{\alpha(1-\alpha)}{2}z^5 + \cdots,$$

respectively. Appealing to (2.4) we see that $k_{\alpha}(z)$ gives the sharp upper bound for $|\gamma_2|$ with any value of α and for $|\gamma_3|$ when $0 \le \alpha \le \frac{2}{3}$, whereas $h_{\alpha}(z)$ provides equality in (2.6) for the remaining values of α .

Theorem 1 shows that no single function serves as the extremal for all coefficients γ_n of inverses for members of \mathcal{S}^*_{α} , $\frac{2}{3} \leq \alpha < 1$, which differs significanty from \mathcal{S} where one function can provide all extremal values. The situa-

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tion for K appears to be surprisingly difficult; (2.8) with $\alpha = \frac{1}{2}$ gives sharp upper bounds $|\gamma_2| \leq 1$ and $|\gamma_3| \leq 1$ when f(z) is in K however $k_1(z)$ cannot give the sharp upper bound for $|\gamma_n|$ for all \mathbb{Z} n. Furthermore it is not likely that using (2.3) and (2.4) and the methods of the theorem can provide the correct bound for γ_4 . However, we can provide an estimate for $|\gamma_n|$.

THEOREM 2. If $F(w) = w + \gamma_2 w^2 + \cdots$ corresponds to f(z) in \mathcal{S}_{∞}^* , then

(2.10)
$$|\gamma_n| \leq \frac{1}{n} \frac{\int (2n(1 - \alpha) + 1)}{\left[\int (n(1 - \alpha) + 1)\right]^2}$$

To establish (2.10) we represent γ_n in a novel way, cf. [5]. Let f(z) and F(w) be as in (1.1) and (1.3) and let c(r) be the image of $||z| = re^{i\Theta} : 0 \le \Theta \le 2$ under f(z), then

$$\gamma_{n} = \frac{1}{2\pi i} \int_{c(r)} \frac{P(w)dw}{w^{n+1}} = \frac{1}{2\pi i} \int_{|z|=r} \frac{zf'(z)}{f(z)^{n+1}} dz =
.11) = (\frac{1}{2\pi i})(\frac{-1}{n}) \left\{ \frac{z}{f(z)^{n}} \Big|_{|z|=r} - \int_{|z|=r} \frac{dz}{f(z)^{n}} \right\} =
= \frac{1}{2\pi i n} \int_{|z|=r} \frac{dz}{f(z)^{n}} \cdot$$

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Now, if f(z) belongs to ω_{∞}^{*} , it is known [4] that $(\frac{z}{f(z)})^{2(1-\alpha)} = 1 + \omega(z)$, where $|\omega(z)| \le |z|$. Consequently, using (2.11) and the principle of subordination We may write

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$$|\gamma_{n}| = \frac{1}{2\pi n} \left| \int_{|z|=r} \left(\frac{z}{f(z)} \right)^{n} \frac{dz}{z^{n}} \right|$$

2.12) $\leq \frac{1}{2\pi nr^{n}} \int_{|z|=r} |1 + \omega(z)|^{2n(1-1)} |dz|$
 $\leq \frac{1}{2\pi nr^{n}} \int_{|z|=r} |1 + z|^{2n(1-\infty)} |dz|.$

Letting $z = re^{1\Theta}$ and replacing r by 1 gives

$$|\gamma_{n}| \leq \frac{2^{2n(1-\alpha)}}{2\pi n} \int_{0}^{2\pi} |\cos \frac{\theta}{2}|^{2n(1-\alpha)} d\theta = \frac{2^{2n(1-\alpha)}}{2\pi n} \int_{0}^{0} (\cos t)^{2n(1-\alpha)} dt = \frac{1}{n} \frac{\int_{0}^{\infty} (2n(1-\alpha)+1)}{[\int_{0}^{\infty} (n(1-\alpha)+1)]^{2}},$$

having made reference to standard tables, [3] for example. For $\alpha = 0$, (2.13) gives $|\gamma_n| \leq \frac{1}{n} {\binom{2n}{n}} = B_n$ which exceeds the correct value given in (1.4). However the orders of both bounds, as $n \rightarrow \infty$, are the same. Also, for $\alpha = 0$, the computations given in (2.12) and (2.13) are equivalent to computing an upper bound for $|\gamma_n|$ when f(z) is the Koebe function (1.2); hence it follows from the work of Baernstein [1] that B_n is an upper bound for coefficients of functions in δ . Of course, this is superfluous in view of Loewner's earlier result, namely (1.4), but it does provide the correct order for $|\gamma_n|$, $n \rightarrow \infty$, with relative case.

It appears then, that bounds for $|\gamma_n|$, f(z) in $\mathcal{S}_{\infty}^{\ell}$ $\alpha \neq 0$, or f(z) in K may be obtainable only with considerable difficulty and that no single member of the class provi-

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des a sharp bound for all indices; on the other hand good bounds for $|a_n|$ are obtainable in a straight forward fashion [2].

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STRESZCZENIE

Otrzymano ostre oszacowania początkowych współczynników dla funkcji odwrotnych do funkcji od-gwiaździstych oraz oszacowania nieostre dla wszystkich współczynników.

Резюме

В работе получено строгие оценки начальных коэффициентов для функций обратных к с -звездным функциям, а также оценки слабие для всех коэффициентов.